

MRX

Magnetic
Reconnection
Experiment

Poloidal Flux Evolution
Null-helicity Reconnection

Princeton Plasma Physics Laboratory, Princeton University

Laboratory experiments available for dedicated study of magnetic reconnection

<i>Device</i>	<i>Where</i>	<i>Since</i>	<i>Geometry</i>	<i>B_g/B_{rec}</i>	<i>Main Issues</i>
3D-CS	Russia	1970	Linear	0-20	Heating
LPD, LAPD	UCLA	1980	Linear	1-10	Heating, waves
TS-3/4	Tokyo	1990	Merging	0-10	Rate, heating
MRX	Princeton	1995	Toroidal, merging	0-20	Rate, heating, waves, boundary, scaling, impulsive
SSX	Swarthmore	1996	Merging	0-1	Heating
VTF	MIT	1998	Toroidal	>~15	Trigger/impulsive, heating
RSX	Los Alamos	2002	Linear	2-40	Boundary, impulsive
RWM	Wisconsin	2002	Linear	10-30	Boundary

Recent Progress in Study of Reconnection Layer and Rate

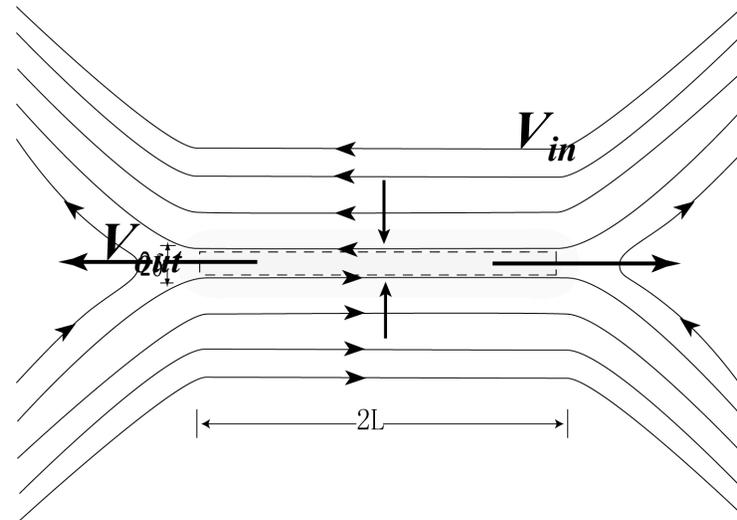
M. Yamada, PPPL

- **Major progress made in**
 - 1) Numerical simulations
 - 2) Space observations
 - 3) Lab experiments
 - Profiles of reconnection layer documented
 - Reconnection rate measured quantitatively
Enhanced resistivity was measured
- **Identified possible causes of fast reconnection**
 - Two fluid physics dominant in the collisionless regime.
Hall effects verified
 - Turbulence/ fluctuations in the sheet
- **Recognized main issues for global reconnection**
 - 3-D MHD modes trigger fast reconnection toroidal pinch systems
 - Impulsive reconnection studied

The Sweet-Parker 2-D MHD Model for Magnetic Reconnection

Assumptions:

- 2D
- Steady-state
- Incompressibility
- Classical Spitzer resistivity



Mass conservation
+ Pressure balance



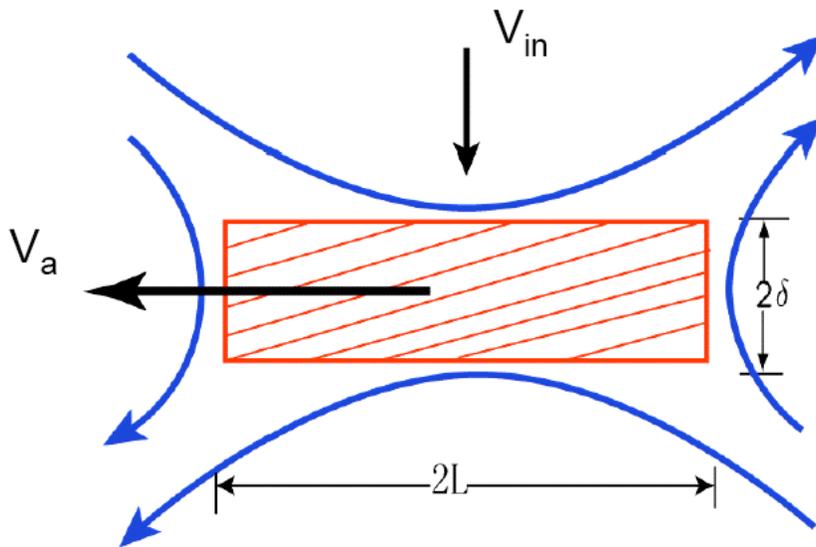
$$\frac{V_{in}}{V_A} = \frac{1}{\sqrt{S}}$$

$$S = \frac{\mu_0 L V_A}{\eta_{Spitz}}$$

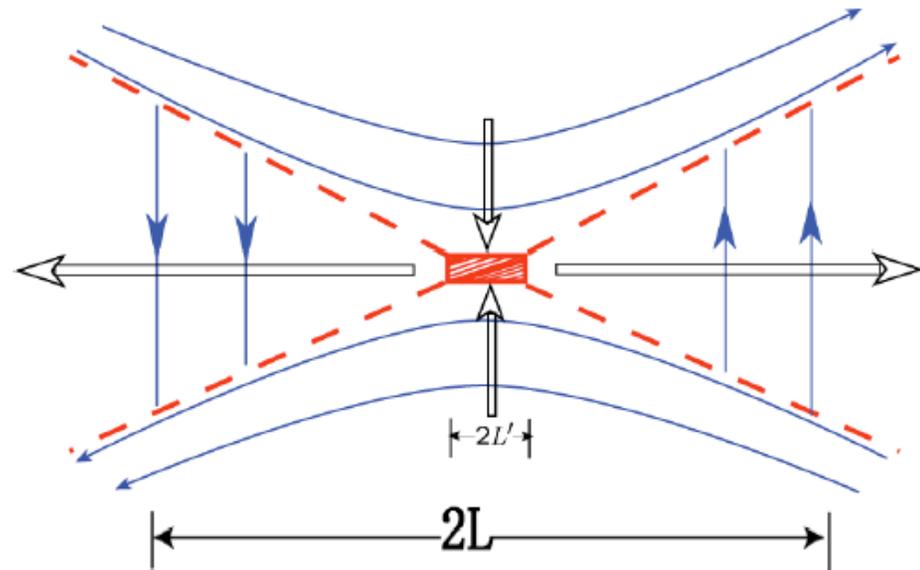
S=Lundquist number

$$\tau_{reconn} \ll \tau_{SP} \sim 3-10 \text{ months}$$

Models for Fast Reconnection



Generalized Sweet-Parker model with **enhanced resistivity caused by turbulence**.



Two-fluid model in which electrons and ions decouple in the diffusion region ($\sim c/\omega_{pi}$)

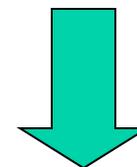
=> Generalized Ohm's law

$$\mathbf{E} + \mathbf{V} \times \mathbf{B} = \eta \mathbf{J} + \frac{\mathbf{J} \times \mathbf{B} - \nabla p}{en} + \frac{m_e}{e^2} \frac{d\mathbf{V}_e}{dt}$$

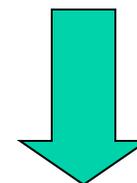
The Hall Effect During Reconnection Shown in 2D Simulation

- **Black lines** → magnetic flux.
- **Blue lines** → ion flow streamlines.
- **Red arrows** → electron flow velocity vectors.
- **Brown arrows** → In-plane current.

Different motions of ions and electrons

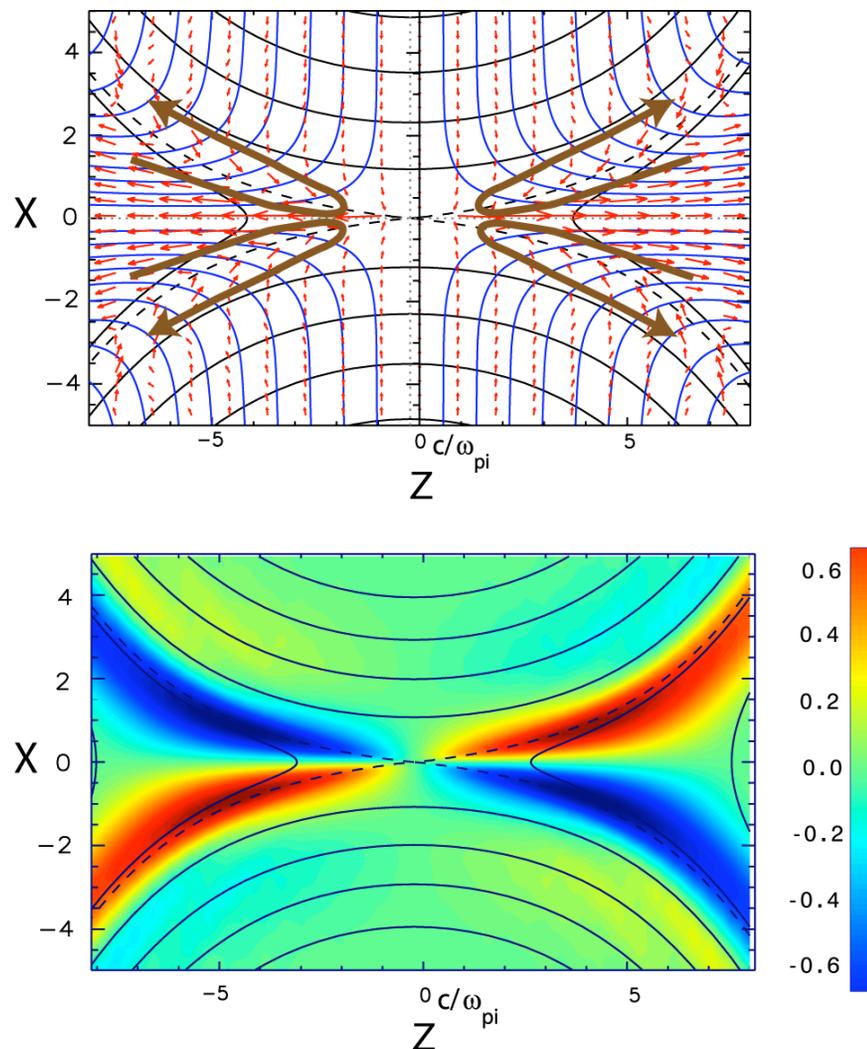


In-plane current

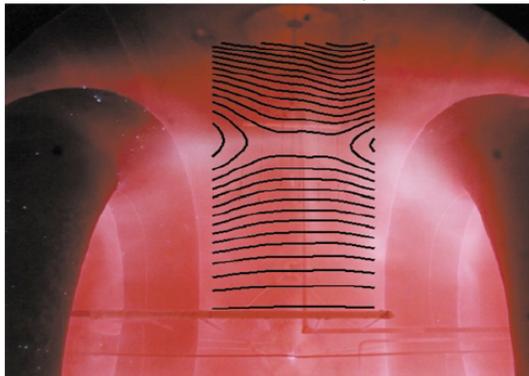
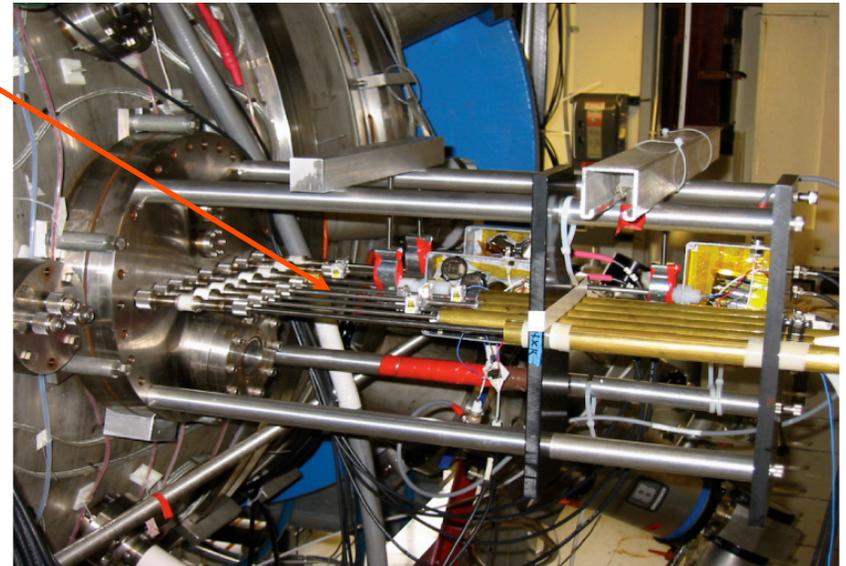
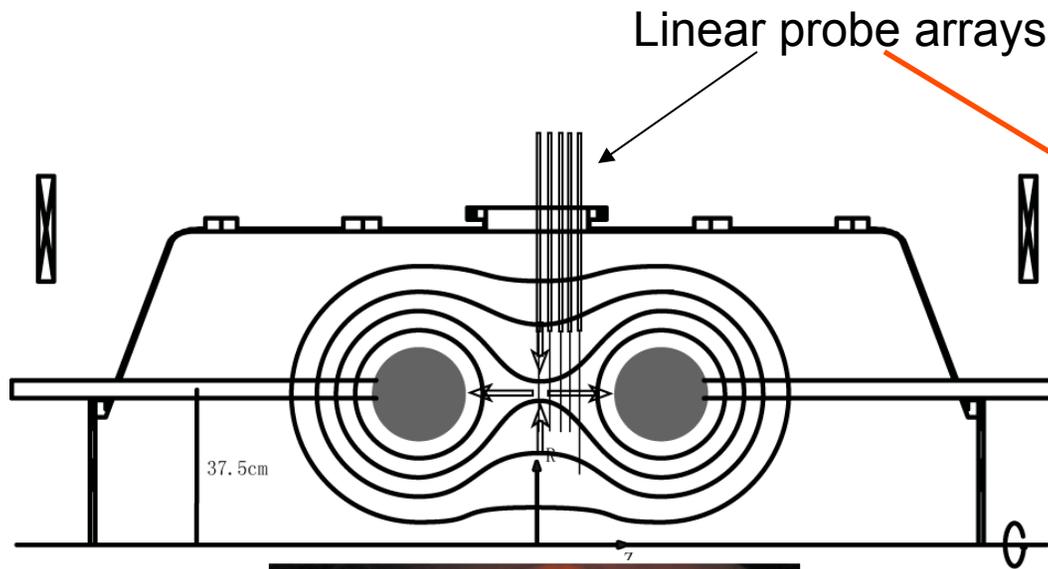


An out-of-plane quadrupole magnetic field

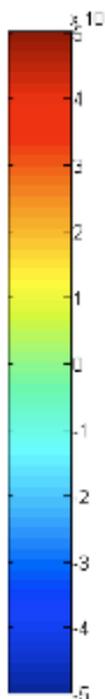
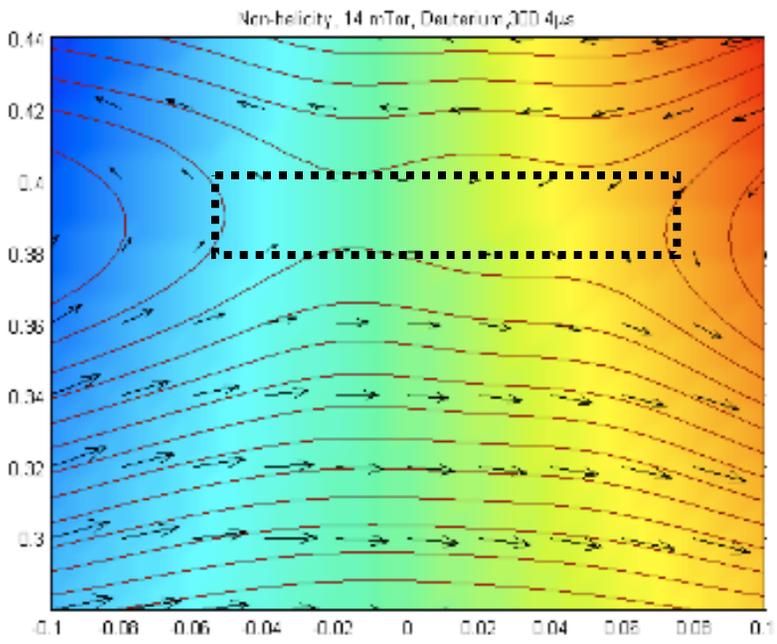
In color code



Dedicated lab experiments systematically studied reconnection layer and rate



- Five fine structure probe arrays with resolution up to $\Delta x = 2.5$ mm in radial direction are placed with separation of $\Delta z = 2-3$ cm



Change of Neutral sheet Profile

from “Rectangular S-P” type
to “Double edge X” shape as
collisionality is reduced

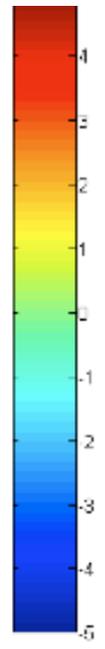
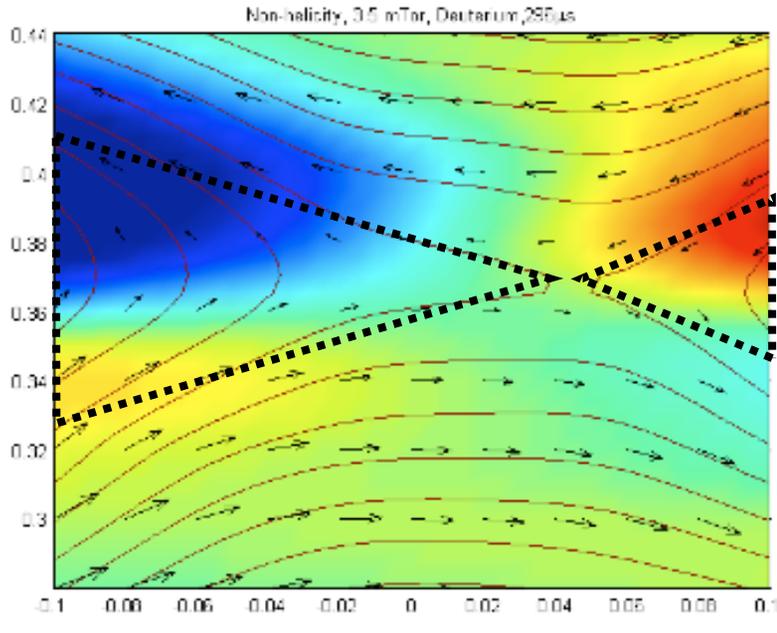
Rectangular shape

Collisional regime: $\frac{\omega}{\nu} \ll \frac{v}{v_{th}}$

$< \frac{\omega}{\nu} \frac{v}{v_{th}}$

Slow reconnection

No Q-P field



X-type shape

Collisionless regime: $\frac{\omega}{\nu} \gg \frac{v}{v_{th}}$

$> \frac{\omega}{\nu} \frac{v}{v_{th}}$

Fast reconnection

Q-P field present

Ma and Bhattacharjee, GRL 1996
Cassak et al PRL 2005

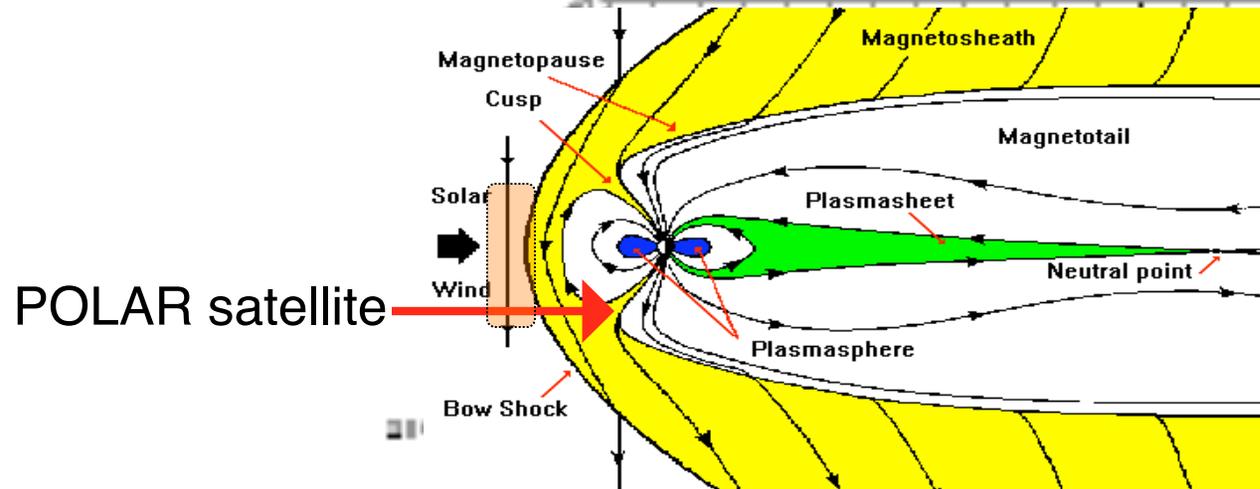
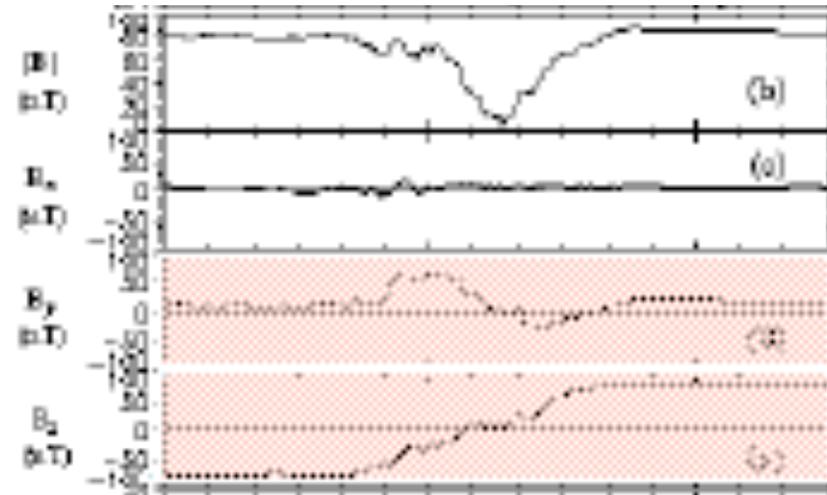
Magnetic Reconnection in the Magnetosphere

A reconnection layer has been documented in the magnetopause

Measurements of Diffusion Region
with a Hall effect signature

Mozer et al., PRL 2002

$$d \sim c/w_{pi}$$



Fast Reconnection \Leftrightarrow Large E_{rec}
 \Rightarrow Enhanced effective resistivity

We have observed:

- **Hall MHD Effects create a large E field (no dissipation)**
- **Electrostatic Turbulence**
- **Electromagnetic Fluctuations**
- **Electron diffusion region identified**

 **How does energy dissipation occur?**

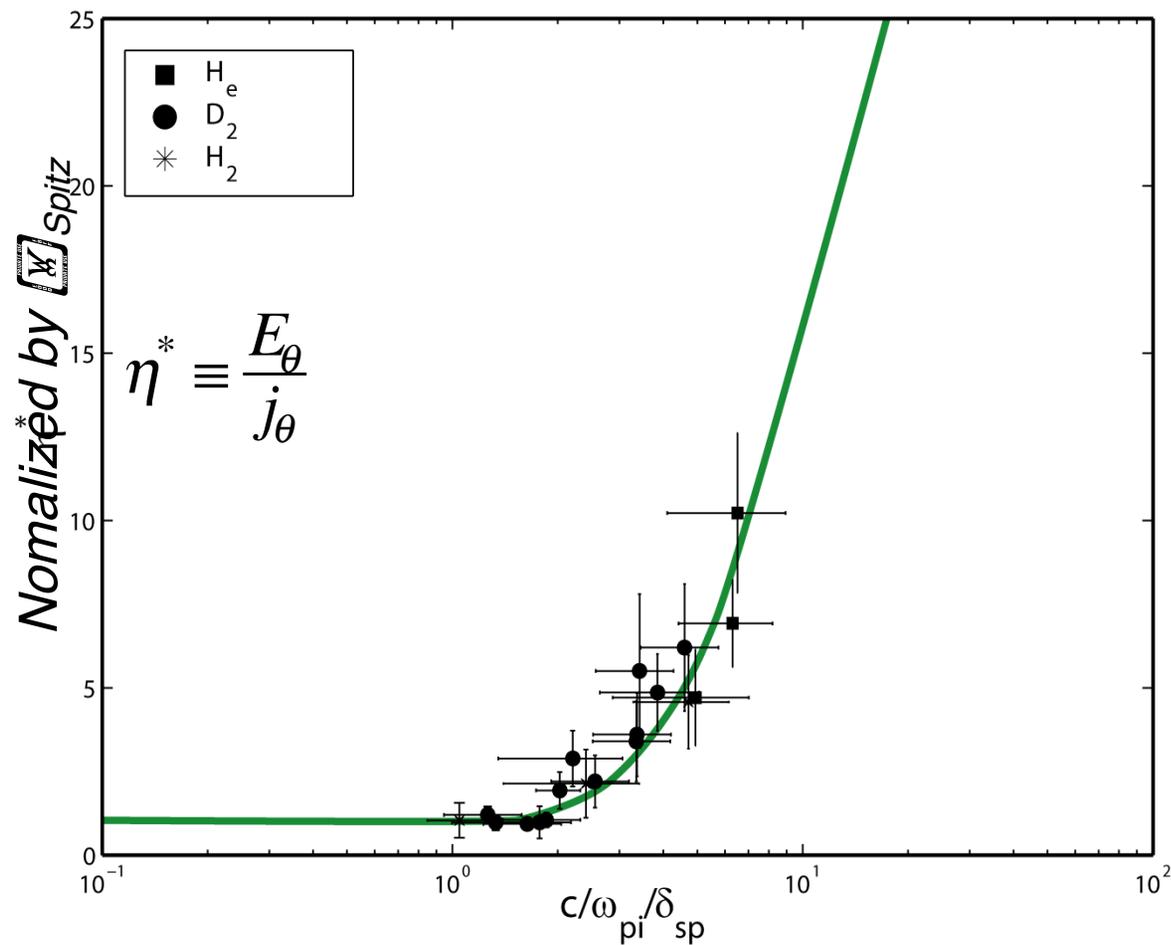
 **How is the reconnection rate really determined?**

What is role of the pressure tensor in Gen-Ohm's law?

\Rightarrow Next level of intensive study

MRX Scaling:

A linkage between space and lab on reconnection



$$\eta_{eff} \sim 20 (\eta_{mfp}/L)$$

$$\sim (c/\omega_{pi})^2 / \eta_{SP}$$

A transition from the MHD to 2 fluid regime when $(c/\omega_{pi}) \sim \eta_{sp}$

Major issues for future reconnection layer research

- **How is the reconnection rate determined?**
 - *Hall effects + Pressure Tensor + Turbulence*
 - *Turbulence*
 - *Multiple reconnection layers*
 - *Impulsive reconnection*
 - *Effects of boundaries*
- **What is a key mechanism for energy conversion (dissipation)?**
=> Particle acceleration and heating
- ***How does it affect global magnetic self-organization?***

Major opportunities for magnetic reconnection research

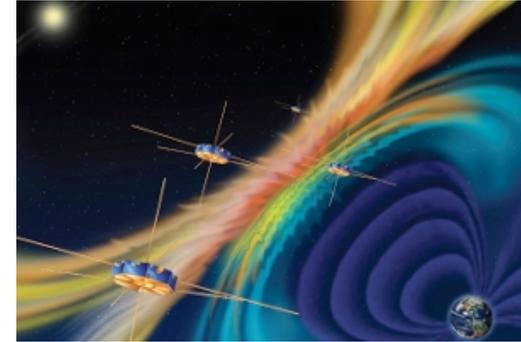
- **Space satellite missions:**

- MMS cluster satellite mission***

- Study multi-scale reconnection regions
with kinetic information*

- Solar satellite missions***

- IRIS, Solar Probe Plus, Chromosphere reconnection, effects of
weakly ionized plasmas, impulsive reconnection*



- **Advanced computer simulations**

- 3-D large S simulation, multiple reconnection layers, kinetic effects*

=> Larger laboratory experiments [\leq NRC 2010 report]

Multiple reconnection layers

Particle heating and acceleration

Impulsive reconnection

Effects of boundary